# Is there ice on the Moon?

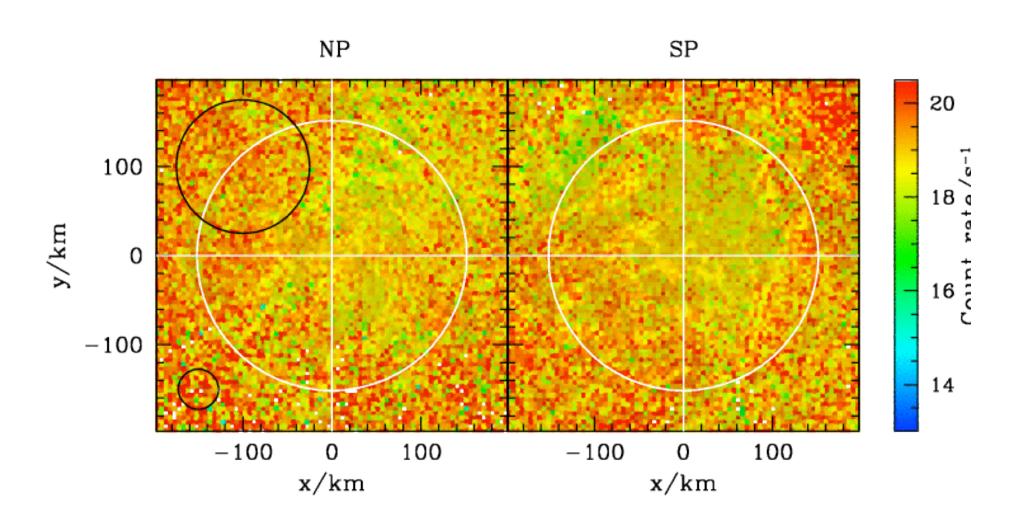
V. Eke (Durham)

L. Teodoro (Glasgow)

R. Elphic (NASA Ames)



# **Lunar Prospector data**



## The problem: D=T\*B+N

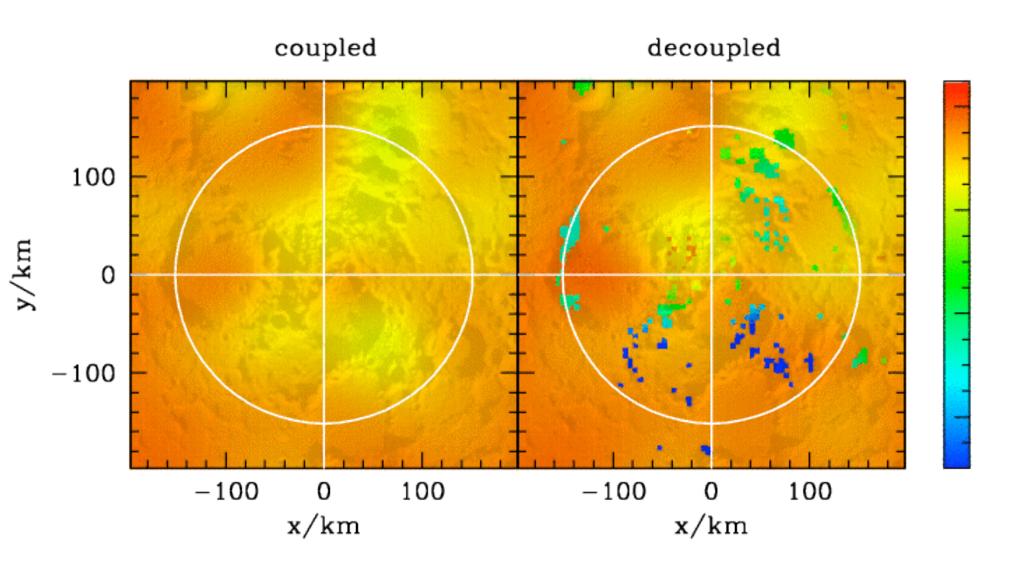
Noise makes the problem ill-posed.

Define 
$$R = D - T_e * B$$
.

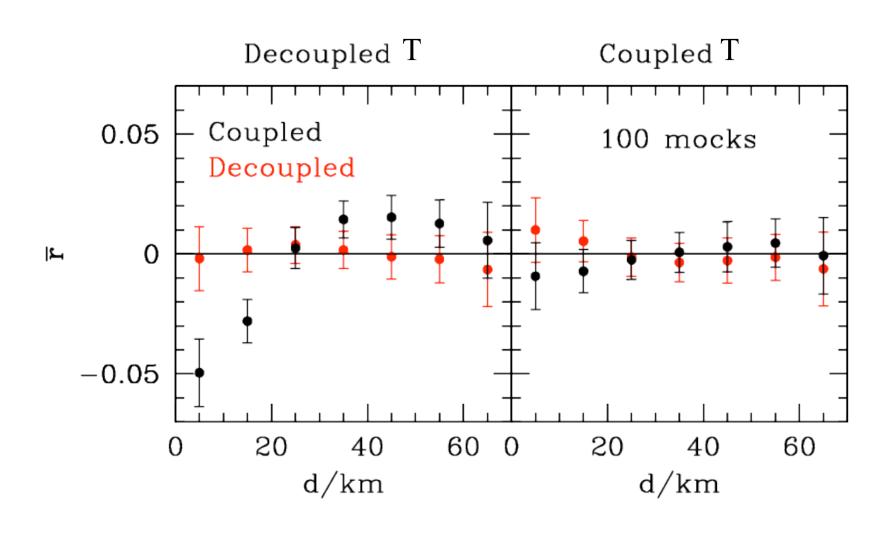
Find a good fit with no details that the data don't demand.

Pixon-based image reconstruction.

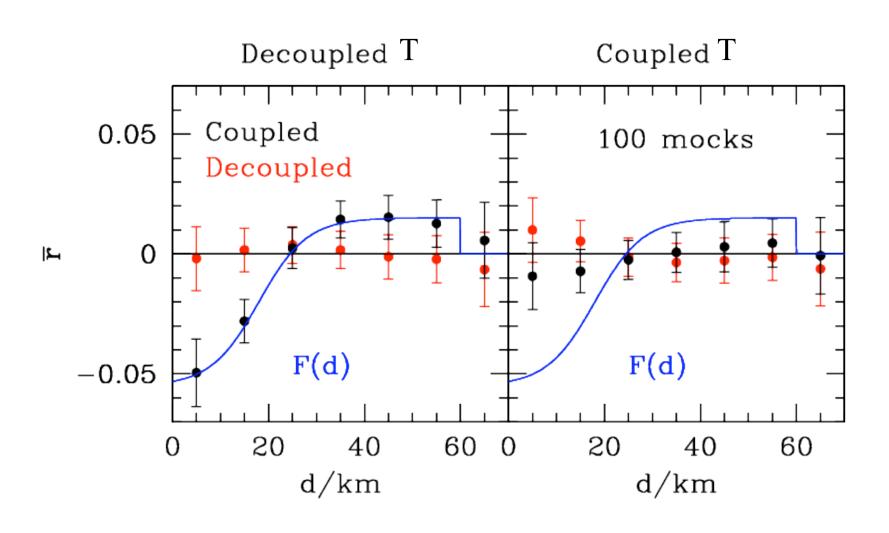
#### **North Pole reconstructions**



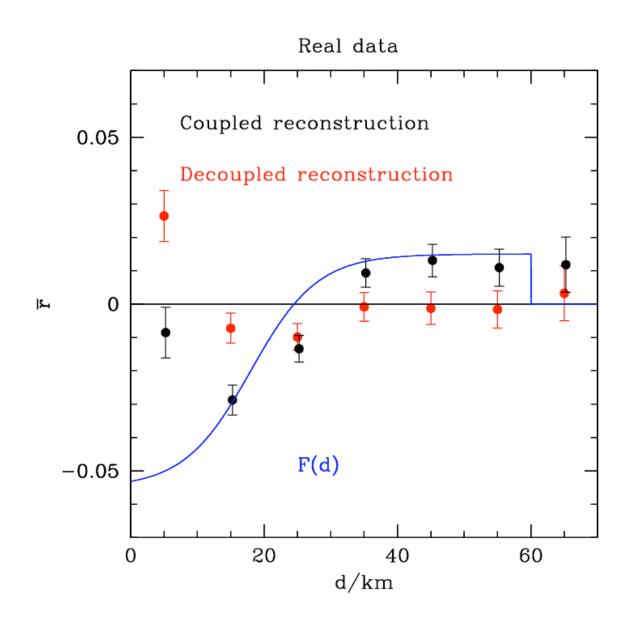
#### Stacked reduced residuals

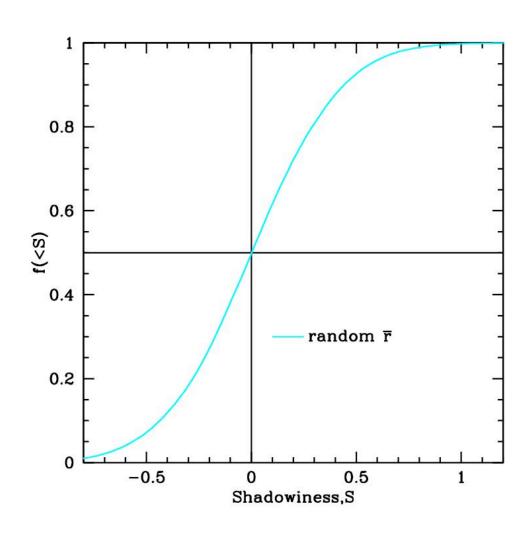


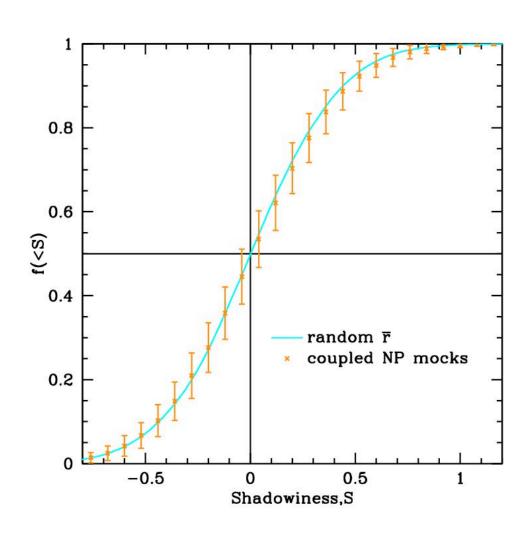
#### Stacked reduced residuals

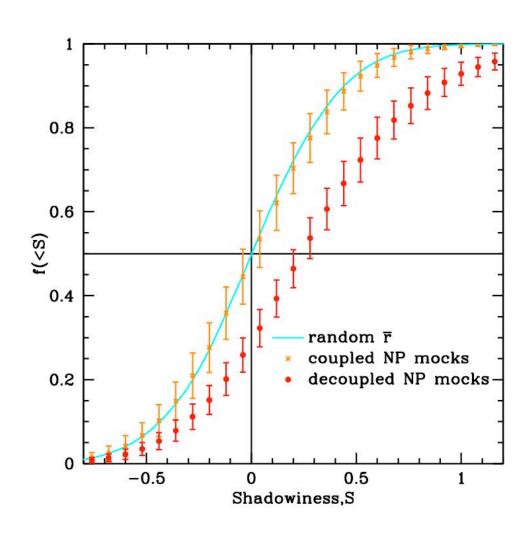


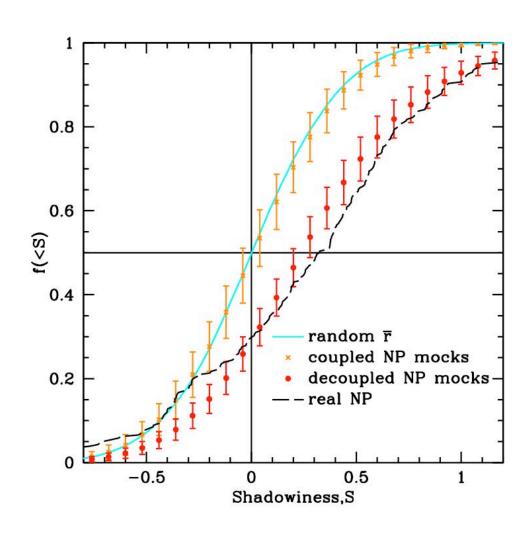
#### Stacked reduced residuals









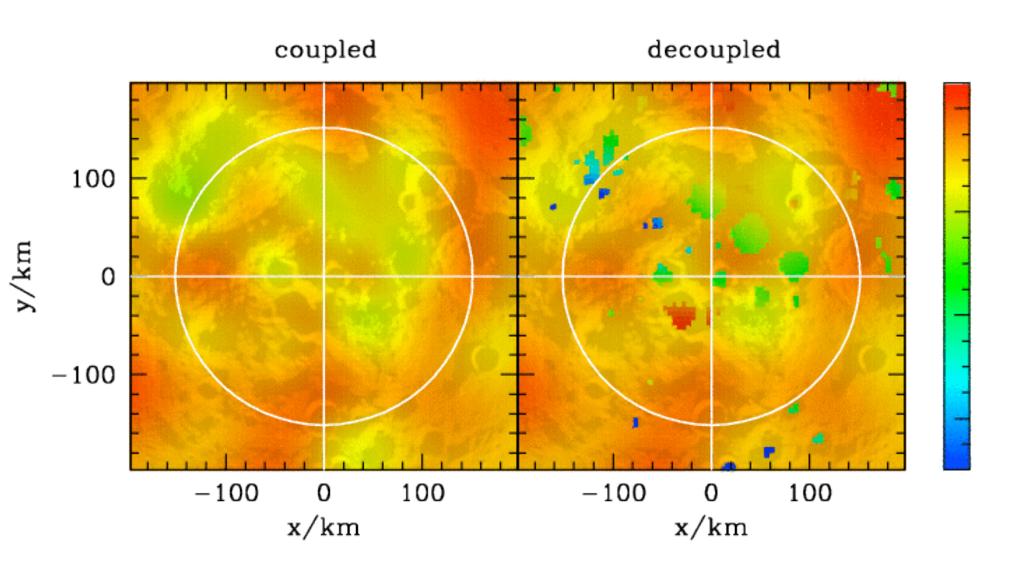


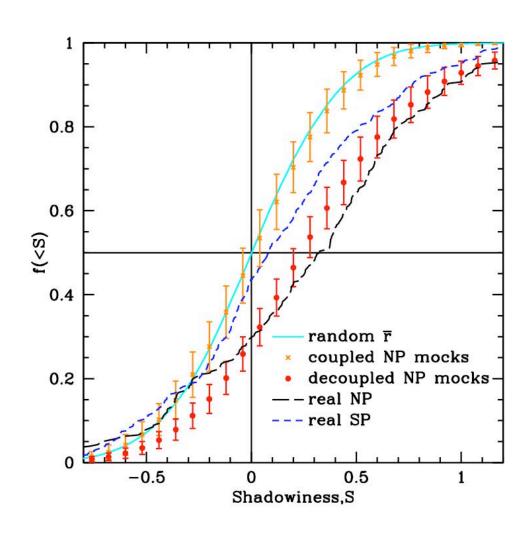
#### Conclusions

- 1) The Lunar Prospector data alone places the excess H into permanently shaded craters.
- 2) The implied concentrations reach ~1% waterequivalent H.
- 3) Including both poles, there is  $\sim 2.10^{11}$ kg H<sub>2</sub>O within 10 degrees of the lunar poles (Lake Havasu holds  $\sim 800$  billion litres).

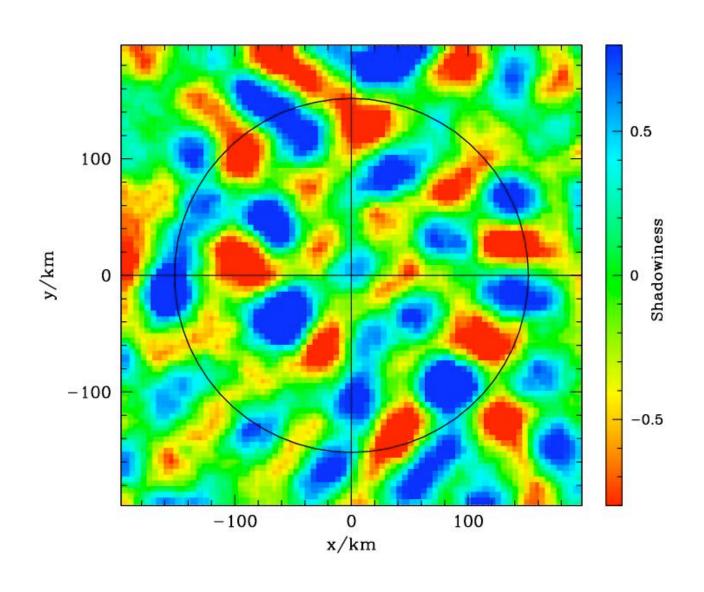


#### South Pole reconstructions

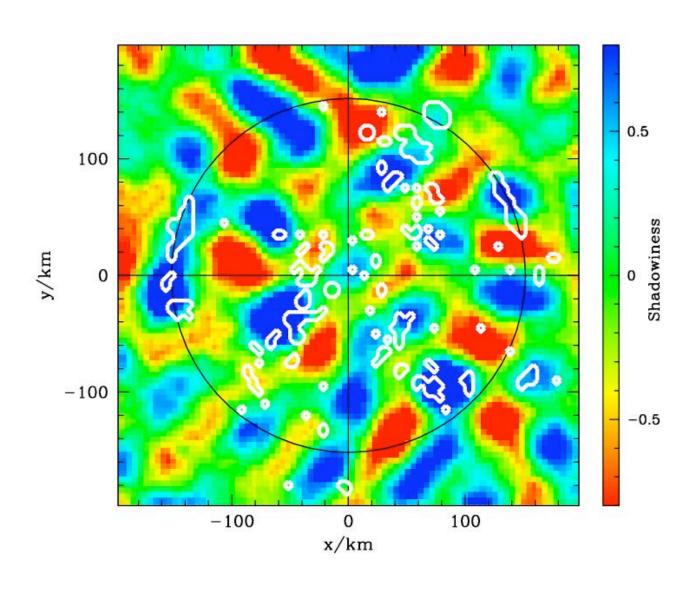




## **NP Shadowiness map**



## **NP Shadowiness map**



### The problem: D=T\*B+N

$$p(D) = 1, p(M) = constant, don't prejudge$$
  
 $\therefore p(T_e \cap M|D) \propto p(D|T_e \cap M)p(T_e|M)$ 

$$p(D|T_e \cap M)$$
: the likelihood

 $p(T_e | M)$  : the image prior

Pixon-based image reconstruction technique